Review of Elementary Quantum Physics

Plank's Radiation Law

• Max Planck (Oct 1900) found formula that reproduced the experimental results

$$\rho(\nu,T) = \frac{8\pi\nu^2}{c^3} < E_{osc} >$$
$$< E_{osc} > = \frac{h\nu}{e^{h\nu/kT} - 1}$$

• derivation from classical thermodynamics, but required assumption that oscillator energies can only take specific values E = 0, hv, 2hv, 3hv, ... (using "Boltzmann factor" $W(E) = e^{-E/kT}$)

Consequences of Planck's hypothesis

oscillator energies E = nhv, n = 0, 1, ...; $h = 6.626 \ 10^{-34} \ Js = 4.13 \ 10^{-15} \ eV \cdot s$ now called Planck's constant

⇒ oscillator's energy can only change by discrete amounts, absorb or emit energy in small packets – "quanta";

 $E_{quantum} = hv$

average energy of oscillator

 $<\!\!E_{osc}\!\!>=h\nu/(e^{x}-1) \text{ with } x=h\nu/kT;$ for low frequencies get classical result $<\!\!E_{osc}\!\!>=kT, k=1.38\cdot 10^{-23} \text{ J}\cdot\text{K}^{-1}$



Metal Foil





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• As blue light strikes the metal foil, the foil emits electrons.

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- When red light hits the metal foil, the foil does not emit electrons.
- Blue light has more energy than red light.
- How could we get more energy into the red light?
- Try increasing the brightness.





- Well, that didn't work!
- Maybe its still not bright enough.





- Still not working.
- What happens with brighter blue light?





• More blue light means more electrons emitted, but that doesn't work with red.





- Wave theory cannot explain these phenomena, as the energy depends on the intensity (brightness)
- According to wave theory bright red light should work!



- Einstein said that light travels in tiny packets called <u>quanta.</u>
- The energy of each quanta is given by its frequency



- Each metal has a minimum energy needed for an electron to be emitted.
- This is known as the work function, W.
- So, for an electron to be emitted, the energy of the photon, *hv*, must be greater than the work function, *W*.
- The excess energy is the kinetic energy, E of the emitted electron.

EINSTEIN'S PHOTOELECTRIC EQUATION:-



The Heisenberg Uncertainty Principle

Before collision: A photon strikes an electron during an attempt to observe the electron's position.





• After collision: The impact changes the electron's momentum, making it uncertain.

The Heisenberg Uncertainty Principle

 If we want accuracy in position, we must use short wavelength photons because the best resolution we can get is about the wavelength of the radiation used. Short wavelength radiation implies high frequency, high energy photons. When these collide with the electrons, they transfer more momentum to the target. If we use longer wavelength, i.e. less energetic photons, we compromise resolution and position.

The Heisenberg Uncertainty Principle

- Symbolically $\Delta x \approx \lambda$
- $\Delta p \approx h/\lambda$
- Δ pλ ≈ h
- $(\Delta p)(\Delta x) \approx h$
- $(\Delta p)(\Delta x) \ge h/2\pi \ge \hbar$

The Heisenberg Uncertainty Principle

- The uncertainty principle can also relate energy and time as follows
- $\Delta t \approx \lambda/c$
- $\Delta E \approx hc/\lambda$
- $(\Delta E)(\Delta t) \approx h$
- $(\Delta E)(\Delta t) \ge h/2\pi \ge \hbar$